SuperB A High-Luminosity Asymmetric e<sup>+</sup>e<sup>-</sup> Super Flavour Factory





Seminar at KEK 29<sup>th</sup> January 2008

FOUNDATION



- SuperB is
  - A Super Flavour Factory with  $L_{peak} > 10^{36}/cm^2/s$
  - An asymmetric energy e<sup>+</sup>e<sup>-</sup> collider
    - Nominal 7 GeV e<sup>-</sup> on 4 GeV e<sup>+</sup> at Y(4S)
    - Flexible running energy & beam polarization options
  - Based on a new approach to collider design
     Avoid limitations due to high beam currents
     (high backgrounds, costly power bill, etc.)
  - The machine to measure new physics flavour couplings in the LHC era



SuperB conceptual design report INFN/AE-07/02, SLAC-R-856, LAL 07-15 (completed April 2007) Available online: http://www.pi.infn.it/SuperB

#### See also

- SuperKEKB Letter of Intent, KEK Report 04-4
- SuperKEKB Physics Working Group, [arXiv:hep-ex/0406071], update in preparation
- J.L.Hewett, D.Hitlin (ed.), SLAC-R-709, [arXiv:hep-ph/0503261]
- Flavour in LHC Era workshops, WG2 report arXiv:0801.1833 [hep-ph]
- "On the Physics Case of a Super Flavour Factory", arXiv:0710:3799 [hep-ph]

### Contents

- Why?
  - Motivation for a Super Flavour Factory in the LHC era
  - Why 10/ab is Not Enough!
- How?
  - Design of SuperB
- Where? When?

## Motivation

- Major challenge for particle physics in the next decade is to go beyond the Standard Model
- Two paths to new physics

1) "relativistic"

New heavy particles produced on mass shell Sensitivity depends on: available centre-of-mass energy knowledge of Standard Model backgrounds

## Motivation

 Major challenge for particle physics in the next decade is to go beyond the Standard Model



### A Tale of Two Frontiers



## LHC-SuperB Interplay

#### 1) LHC discovers new physics

- Can it be flavour blind? (ie. no signals in flavour)
  - No, it must couple to SM, which violates flavour
  - Any TeV scale NP model includes new flavoured particles
- What is the minimal flavour violation? (ie. worst case)
  - NP follows SM pattern of flavour and CP violation
  - Super*B* detects NP effects for particle masses up to >600 GeV
- What if NP flavour couplings are not suppressed?
  - SuperB measures NP flavour couplings and distinguishes models

#### 2) LHC does not discover new physics

- Problem for naturalness?
  - Not really just an order of magnitude argument
- How to probe higher mass scales?
  - NP models with unsuppressed flavour couplings can reach scales of 10s, 100s or even 1000s of TeV

## LHC-SuperB Interplay

- Flavour observables are complementary to those at the energy frontier
  - measure different new physics parameters
  - powerful to distinguish models



Flavour Observables Sensitive to New Physics  $\Delta m_{\kappa} \epsilon_{\kappa} \epsilon_{\kappa} \delta(K_{I} \to \pi^{0} \nu \overline{\nu}) B(K^{+} \to \pi^{+} \nu \overline{\nu}) B(K^{+} \to I^{+} \nu)$  $\Delta m_d \quad A_{SI}(B_d) \quad S(B_d \rightarrow J/\psi K_S) \quad S(B_d \rightarrow \phi K_S)$  $\alpha(B \rightarrow \pi \pi, \rho \pi, \rho \rho) \qquad \gamma(B \rightarrow DK) \qquad CKM \text{ fits}$  $\Delta m_{s} \quad A_{s}(B_{s}) \quad S(B_{s} \rightarrow J/\psi\phi) \quad S(B_{s} \rightarrow \phi\phi)$  $B(b \rightarrow s \gamma) \quad A_{CP}(b \rightarrow s \gamma) \quad S(B^{0} \rightarrow K_{S} \pi^{0} \gamma) \quad S(B_{s} \rightarrow \phi \gamma)$  $B(b \rightarrow d\gamma) \quad A_{CP}(b \rightarrow d\gamma) \quad A_{CP}(b \rightarrow (d+s)\gamma) \quad S(B^{0} \rightarrow \rho^{0}\gamma)$  $B(b \rightarrow s I^+ I^-) \quad B(b \rightarrow d I^+ I^-) \quad A_{FB}(b \rightarrow s I^+ I^-) \quad B(b \rightarrow s v \overline{v})$  $B(B_{c} \rightarrow I^{+}I^{-}) \quad B(B_{d} \rightarrow I^{+}I^{-}) \quad B(B^{+} \rightarrow I^{+}\nu)$  $B(\mu \rightarrow e \gamma) \quad B(\mu \rightarrow e^+ e^- e^+) \quad (g-2)_\mu \quad \mu \quad EDM$  $B(\tau \rightarrow \mu \gamma) \quad B(\tau \rightarrow e \gamma) \quad B(\tau^+ \rightarrow I^+ I^- I^+) \quad \tau \quad CPV \quad \tau \quad EDM$  $B(D_{(s)}^+ \rightarrow I^+ v)$   $X_D Y_D$  charm CPV 10 ... add your favourite here ...

## Good News and Bad News

#### Bad news

- no single "golden mode"
- (of course, some channels preferred in certain models)
- Good news
  - multitude of new physics sensitive observables
  - maximize sensitivity by combining information
  - correlations between results distinguish models

SuperB "treasure chest" of new physics sensitive flavour observables





#### Focus on theoretically clean channels

no theory improvements needed	β(J/ψ K), γ(DK), α(ππ)*, lepton FV and UV, S(ρ <sup>0</sup> γ) CPV in B->Xγ, D and τ decays zero of FB asymmetry B->X <sub>s</sub>   <sup>+</sup>   <sup>-</sup>	NP insensitive or null tests of the SM or SM already known with the required accuracy
improved lattice QCD	meson mixing , B->D(*)lv,B->π(ρ)lv, B->K*γ, B->ργ, B->lv, B <sub>s</sub> ->μμ	target error: ~1-2% Feasible
improved OPE+HQE	B->X <sub>u,c</sub> Iv, B->Xγ	target error: ~1-2% Possibly feasible with SuperB data getting rid of the shape function. Detailed studies required
improved QCDF or SCET or flavour symmetries	S's from TD A <sub>æ</sub> in b -> s transitions	target error: ~2-3% large and hard to improve uncertainties on small corrections. In addition, FS+data can bound the theoretical error

DECREASING THEORETICAL UNCERTAINTY

table by M.Ciuchini

## Super Flavour Factory

- Data taken at Y(4S) allows studies of B, tau, charm, charmonia, ISR, γγ physics (and more)
- Super*B* is designed with flexible running energy
  - charm-tau threshold region
  - other Upsilon resonances including Y(5S)  $\Rightarrow can study B_s sector, including \Delta\Gamma_s and \varphi_s (but not \Delta m_s)$
- Considering beam polarization option
  - study chiral structure of various processes
  - significant improvement in sensitivity for  $\tau$  EDM

see arXiv:0707.1658 and arXiv:0707.2496 14

## Lepton Flavour Violation

 Observable LFV signals predicted in a wide range of models, including those inspired by Majorana neutrinos



#### **Much** better sensitivity than LHC, even for $\tau \rightarrow \mu \mu \mu$

#### Lepton Flavour & Neutrino Physics

 In many scenarios, LFV rates are linked to (both low and high energy) neutrino parameters



### Charm at SuperB

SuperB can study the full range of charm phenomena – including CP violation

HFAG-charm

FPCP 2007

**\_0.04** 

0.03

0.02

0.01

0

CP violation in charm highly sensitive new physics probe

sensitivity:  $\phi_{D} \sim 1^{\circ}$ 

Mode	Observable	$B$ Factories (2 $\rm ab^{-1})$	$SuperB$ (75 $ab^{-1}$ )
$D^0 \to K^+ K^-$	УСР	$2 - 3 \times 10^{-3}$	$5 \times 10^{-4}$
$D^0 \to K^+ \pi^-$	$y'_D$	$23 \times 10^{-3}$	$7 \times 10^{-4}$
	$x_D^{\prime 2}$	$1-2 \times 10^{-4}$	$3 \times 10^{-5}$
$D^0 \to K^0_{\scriptscriptstyle S} \pi^+ \pi^-$	$y_D$	$2 - 3 \times 10^{-3}$	$5 \times 10^{-4}$
	$x_D$	$23 \times 10^{-3}$	$5 \times 10^{-4}$
Average	$y_D$	$1-2 \times 10^{-3}$	$3 \times 10^{-4}$
	$x_D$	$2 - 3 \times 10^{-3}$	$5 \times 10^{-4}$







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#### Leptonic B Decays



#### Forward-Backward Asymmetry

Exclusive: B→K\*II

Inclusive: b→sll



Inclusive is much cleaner  $\Leftrightarrow$  need SuperB statistics

#### SuperB UT fit scenarios



"the dream"



• Possible NP discovery from precise CKM metrology

• Precise knowledge of SM parameters essential in any scenario 20



By M.Ciuchini

#### SuperB physics in tables

Observable	$B$ factories (2 $ab^{-1}$ )	Super $B$ (75 ab <sup>-1</sup> )
$sin(2\beta) (J/\psi K^0)$	0.018	0.005 (†)
$cos(2\beta) (J/\psi K^{*0})$	0.30	0.05
$sin(2\beta)$ (Dh <sup>0</sup> )	0.10	0.02
$cos(2\beta)$ (Dh <sup>0</sup> )	0.20	0.04
$S(J/\psi \pi^0)$	0.10	0.02
$S(D^{+}D^{-})$	0.20	0.03
$S(\phi K^0)$	0.13	0.02 (*)
$S(\eta' K^0)$	0.05	0.01 (*)
$S(K_{S}^{0}K_{S}^{0}K_{S}^{0})$	0.15	0.02 (*)
$S(K_{S}^{0}\pi^{0})$	0.15	0.02 (*)
$S(\omega \tilde{K}_{s}^{0})$	0.17	0.03 (*)
$S(f_0 \tilde{K}_{S}^{0})$	0.12	0.02 (*)
$\gamma (B \rightarrow DK, D \rightarrow CP \text{ eigenstat})$	$\sim 15^{\circ}$	2.5 °
$\gamma (B \rightarrow DK, D \rightarrow \text{suppressed st})$	ates) $\sim 12^{\circ}$	2.0°
$\gamma (B \rightarrow DK, D \rightarrow \text{multibody st.}$	ates) $\sim 9^{\circ}$	1.5 °
$\gamma (B \rightarrow DK, \text{ combined})$	$\sim 6^{\circ}$	1-20
$\alpha (B \rightarrow \pi \pi)$	$\sim 16^{\circ}$	30
$\alpha (B \rightarrow \rho \rho)$	$\sim 7^{\circ}$	$1-2^{\circ}$ (+)
$\alpha (B \rightarrow \rho \pi)$	$\sim 12^{\circ}$	2°
a (combined)	$\sim 6^{\circ}$	1-2° (+)
$2\beta + \gamma (D^{(*)\pm}\pi^{\mp}, D^{\pm}K^{0}_{S}\pi^{\mp})$	20°	5°
V <sub>cb</sub> (exclusive)	4% (*)	1.0% (*)
V <sub>cb</sub> (inclusive)	1% (+)	0.5% (*)
$ V_{ub} $ (exclusive)	8% (+)	3.0% (*)
V <sub>ub</sub> (inclusive)	8% (+)	2.0% (*)
$BR(B \rightarrow \tau \nu)$	20%	4% (†)
$BR(B \rightarrow \mu\nu)$	visible	5%
$BR(B \rightarrow D\tau\nu)$	10%	2%
$BR(B \rightarrow \rho\gamma)$	15%	3% (†)
$BR(B \rightarrow \omega \gamma)$	30%	5%
$A_{CP}(B \rightarrow K^* \gamma)$	0.007 (†)	0.004 († +)
$A_{CP}(B \rightarrow \rho \gamma)$	$\sim 0.20$	0.05
$A_{CP}(b \rightarrow s\gamma)$	0.012 (†)	0.004 (†)
$A_{CP}(b \rightarrow (s + d)\gamma)$	0.03	0.006 (†)
$S(K_e^0 \pi^0 \gamma)$	0.15	0.02 (*)
$S(\rho^{0}\gamma)$	possible	0.10
$A_{CP}(B \rightarrow K^* \ell \ell)$	7%	1%
$A^{FB}(B \rightarrow K^* \ell \ell) s_0$	25%	9%
$A^{FB}(B \rightarrow X_s \ell \ell) s_0$	35%	5%
$BR(B \rightarrow K\nu\overline{\nu})$	visible	20%
$BR(B \rightarrow \pi \nu \bar{\nu})$	-	possible

Mode	Observable $B \to F$	actories (2 ab	$^{-1}$ ) SuperB (75 ab <sup>-1</sup> )	
$D^0 \to K^+ K^-$	9 <sub>CP</sub>	$23 \times 10^{3}$	$5 \times 10^{-4}$	
$D^0 \to K^+\pi^-$	$y'_D$	$23\times10^{3}$	$7 \times 10^{-4}$	
	$x'^{2}_{D}$	$12\times10^{-4}$	$3 \times 10^{-5}$	
$D^0 \to K^0_s \pi^+ \pi^-$	$y_D$	$23\times10^{3}$	$5 \times 10^{-4}$	
	$x_D$	$23\times10^{1}$	$5 \times 10^{-4}$	
Average	УD	$12\times10^{-3}$	$3 \times 10^{-4}$	
	$x_D$	$23 \times 10^{1}$	$5 \times 10^{-4}$	Sensitivity
5	10.4	$D^0$	$\rightarrow e^+e^-, D^0 \rightarrow \mu^+\mu^-$	$1 \times 10^{-8}$
5	-10X	$D^0$	$\rightarrow \pi^0 e^+ e^-$ , $D^0 \rightarrow \pi^0 \mu^+ \mu^-$	$2 \times 10^{-8}$
		$D^0$	$\rightarrow \eta e^+ e^-, D^0 \rightarrow \eta \mu^+ \mu^-$	$3 \times 10^{-8}$
Impro	ovemen	$D^0$	$\rightarrow K_s^0 e^+ e^-, D^0 \rightarrow K_s^0 \mu^+ \mu$	$i^{-}$ 3 × 10 <sup>-8</sup>
_		$D^+$	$\rightarrow \pi^+ e^+ e^-,  D^+ \rightarrow \pi^+ \mu^+ \mu^-$	$\mu^{-}$ 1 × 10 <sup>-8</sup>
Process	Sensitivity	7		
$B(\tau \rightarrow \mu)$	$\gamma$ ) 2 × 10 <sup>-9</sup>	- D <sup>0</sup>	$\rightarrow e^{\pm}\mu^{\mp}$	$1 \times 10^{-8}$
$B(\tau \rightarrow cc)$	$2 \times 10^{-9}$	$D^+$	$\rightarrow \pi^+ e^{\pm} \mu^{\mp}$	$1 \times 10^{-8}$
$B(\tau \rightarrow e\gamma)$	y) 2 × 10 -	$D^0$	$\rightarrow \pi^0 e^{\pm} \mu^{\mp}$	$2 \times 10^{-8}$
$B(\tau \rightarrow \mu \mu)$	$(\mu \mu) = 2 \times 10^{-10}$	$D^0$	$\rightarrow \eta e^{\pm} \mu^{\mp}$	$3 \times 10^{-8}$
$\mathcal{B}(\tau \rightarrow ee$	(e) $2 \times 10^{-10}$	$D^0$	$\rightarrow K_s^0 e^{\pm} \mu^{\mp}$	$3 \times 10^{-8}$
$\mathcal{B}(\tau \rightarrow \mu \eta)$	$4 \times 10^{-10}$			
$\mathcal{B}(\tau \rightarrow e\eta)$	$6 \times 10^{-10}$	D+	$\rightarrow \pi^- e^+ e^+,  D^+ \rightarrow K^- e^+$	$e^+$ 1 × 10 <sup>-8</sup>
$B(\tau \rightarrow \ell k)$	$\binom{0}{2}$ 2 × 10 <sup>-10</sup>	D+	$\rightarrow \pi^- \mu^+ \mu^+, D^+ \rightarrow K^- \mu^+$	$^{+}\mu^{+} = 1 \times 10^{-8}$
	·s, 2 / 10	— D <sup>+</sup>	$\rightarrow \pi^- e^\pm \mu^\mp,  D^+ \rightarrow K^- e^\pm$	$\mu^{\mp} = 1 \times 10^{-8}$
+τFC ph	ysics (CPV,	)		
Super Fl	lavour Fac	tory	Observable	Error with 1 $ab^{-1}$
			$\Delta\Gamma$	$0.16 \ ps^{-1}$
at	reasure ch	iest"	Г	$0.07 \text{ ps}^{-1}$
	of	new	$\beta_{s}$ from angular analysis	20 <sup>n</sup>
- Constant			$A_{SL}^{*}$	0.006
10000	phys	SICS-	A <sub>CH</sub>	0.004
N A	sens	itive	$\mathcal{B}(B_s \rightarrow \mu^+ \mu^-)$	-
02	Joing		$B(R \rightarrow \infty)$	38%
19	observo	ibles	β. from J/ψφ	10°
No.			A CONTRACT OF A CONTRACT.	

M.Ciuchini at SuperB Review, LNF, 12 November

# Why 10/ab Is Not Enough! Just a few examples ...

- Lepton flavour violation
  - Need a big push into the unexplored region
- Forward-backward asymmetry in  $b \rightarrow sll$ 
  - Must improve beyond 10% theory error of exclusive modes
- Rare B decays  $(B \rightarrow K^{(*)}vv, B \rightarrow \mu^{+}\mu^{-})$

- Prospects for observation marginal at 10/ab

 Null tests, such as CP violation in charm Limited only by statistics

## SuperB: How?

- Physics case for Super Flavour Factory is compelling
- Luminosity should be above 10<sup>36</sup>/cm<sup>2</sup>/s
  - Enables integration of over 10/ab/year
  - Backgrounds and running efficiency should be comparable to current B factories
  - Power consumption should be affordable
- Attempts to upgrade PEP-II and KEKB with high current hit limitations due to beam instabilities, backgrounds and power

### How Can it be Achieved?

Luminosity must be  $\sim 10^{36}$ /cm<sup>2</sup>/s or higher

- Enables integration of over 10/ab/year
- Two orders of magnitude higher than now
- $\Rightarrow$  Push current B factories to the limit (SuperKEKB)



### Alternative approach

- Machine is based on ILC damping ring lattice
  - High luminosity through small emittance
  - Collide with large Piwinski angle & crab waist



### **Collision Scheme**

Large Piwinski angle ( $\varphi = \theta \sigma_z / \sigma_x$ ) & "crab waist"



- High luminosity, Low currents, Small backgrounds
- Stable dynamic aperture
- → Wall plug power ~ 20 MW

#### SuperB Parameters

L (cm <sup>-2</sup> s <sup>-1</sup> )	<b>1.x10</b> <sup>36</sup>	28
Power (MW)	17	ingrici iuriniosides:
σ <sub>z</sub> (mm)	5	higher luminosities
σ <sub>x</sub> * (μm) (LER/HER)	10/6	Future upgrades
σ <sub>y</sub> * (μm) (LER/HER)	0.039	
$\beta_{x}^{*}$ (mm) (LER/HER)	35/20	
$\beta_{y}^{*}$ (mm) (LER/HER)	0.22/0.39	
$ε_y$ (pm-rad) (LER/HER)	7/4	
$ε_x$ (nm-rad) (LER/HER)	2.8/1.6	
θ (rad)	2x24	
No. part/bunches	5.5x10 <sup>10</sup>	
No. bunches	1342	
Current (A)/beam	2.	
Energy (GeV) (LER/HER)	4/7	
Circumference (m)	1800.	

## Crab Waist Beam Tests

- Tests ongoing at DaΦNe accelerator (LNF)
- Also planned as part of DaΦNe upgrade
   Will reach peak luminosity of 10<sup>33</sup>/cm<sup>2</sup>/s
- Details in talk of M.Biagini at BNM2008 http://superb.kek.jp/bnm2008/slide/3-Biagini.ppt
- Results promising so far ...

#### Half IR1 Magnetic Layout



### Good News

- Although collider scheme is completely new, it can be constructed largely by recycling existing hardware (eg. PEP-II magnets)
- Some new hardware required (eg. sextupole magnets for crab waist)
- Backgrounds comparable to current B factories, so SuperB detector can be based on BaBar (or Belle)

#### Significant cost savings!

## Backgrounds

- Dominated by QED cross section
  - Low currents / high luminosity
    - Beam-gas are not a problem
    - SR fan can be shielded

	Cross section	Evt/bunch xing	Rate	
Radiative Bhabha	~340 mbarn ( Eγ/Ebeam > 1% )	~680	0.3THz	
e⁺e⁻ pair production	~7.3 mbarn	~15	7GHz	<i>p</i>
Elastic Bhabha	O(10 <sup>-5</sup> ) mbarn (Det. acceptance)	~20/Million	10KHz	
Υ(4S)	O(10 <sup>-6</sup> ) mbarn	~2/million	I KHz	$p_+$

## Backgrounds and Detectors

- Backgrounds depend on various factors
  - luminosity
    - radiative BhaBha scattering
    - e<sup>+</sup>e<sup>-</sup> pair production
  - currents
    - synchrotron radiation
    - beam-gas interaction
  - beam size
    - Touschek scattering
    - beam-beam interactions

luminosity lifetime ~ 5 minutes

main problem for SuperKEKB: beam backgrounds ~ 20 x today

possible problem for SuperB: motivates smaller beam asymmetry (7 GeV on 4 GeV)

- Interaction point design & shielding requires care
- Detector can be based on existing BaBar / Belle

#### Interaction Region Design



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## Detector R&D

- Detector R&D required for most subsystems
  - vertex detector
    - first layer close (~1cm) to beam spot
    - use pixels or striplets to cope with occupancy
  - particle identification
    - improved readout for barrel (DIRC)

improvements in hermeticity important for many measurements

- forward PID device under consideration
- calorimeter
  - LYSO based forward endcap
  - backward endcap (veto counter?) under consideration
- electronics, trigger, DAQ & offline computing
  - need to deal with high physics trigger rate



## SuperB: Where and When?

- Physics case for Super Flavour Factory is compelling
- Machine and detector can be realised at reasonable cost
- How can this project fit into global and regional roadmaps?
  - Well established priorities: LHC, ILC, neutrinos
  - The physics potential right timescale (2010s) European Strategy Strategy Grouping (CERN Council Strategy Group) Scope for additional regional project

#### Potential SuperB site on the University of Rome Tor Vergata campus



#### Potential SuperB site on the University of Rome Tor Vergata campus



#### Footprint



by F.Forti

#### **Recent Progress and Future Plans**

- CDR is being evaluated by an International Review Committee (IRC)
- Continuing work on
  - Physics case
  - Detector R&D
  - Accelerator design
- Expect IRC report April 2008
  - If positive, will request endorsement from CERN strategy group
  - After this milestone, will request funding

#### International Review Committee

- R. Petronzio, President of INFN, formed an International Review Committee to evaluate the SuperB CDR
- The committee members are:

   J. Dainton (chair) [UK]
   J. LeFrancois [France]
   H. Aibara [Japan]
   P. Houer [Cormany]
   K.K.
  - H. Aihara [Japan]R. Heuer [Germany]Y.-K. Kim [US]A. Masiero [Italy]A. Seiden [US]D. Shulte [CERN]
- Meeting with the committee held November 2007
- Requests for further information being responded to
- Expect final report April 2008

#### **Ongoing Activity**



#### SuperB Workshop VI

New Physics at the Super Flavour Factory SuperB

#### IFIC, Valencia, 7-15 January, 2008

Goals

Sharpening the discovery potential of the Super Flavour Factory
 Simulation studies including detector response and machine parameters

#### Programme Committee:

D. Asner (U. Carleton)

- M. Ciuchini (INFN, Rome-III)
- R. Faccini (INFN, Rome-I)
- M. Giorgi (INFN, Pisa) Chair
- D. Hitlin (Caltech)
- D. Hitim (Caneen)
- J. Olsen (*U. Princeton*)
- M. Roney (U. Victoria)
- A. Stocchi (LAL-Orsay)

Local Organizing Committee:

- J. Bernabéu (*IFIC*) *Chair* F. Martínez-Vidal (*IFIC*) A. Oyanguren (*IFIC*)
- M.A. Sanchis-Lozano (1FIC)
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#### SuperB Detector I

#### A workshop on design and R&D for a SuperB detector

The workshop will focus on the detector for the high luminosity 10<sup>36</sup> cm<sup>-2</sup>s<sup>-1</sup> SuperB flavor factory. Particular emphasis will be placed on the R&D needed for detector subsystems, and on the development of a suite of software tools needed for simulation studies.

http://www-conf.slac.stanford.edu/superB2008/

SuperB URL: http://www.pi.infn.it/SuperB/

Contact information: Thanh Ly tkl@slac.stanford.edu 650-926-4496

Rafel Montaner, Valencia Cincuenta científicos de la UE, EE UU v Canadá estarán reunidos en Valencia hasta el martes con el fin de concluir el diseño de un supercolisionador de partículas subatómicas único en el mundo que aportará nuevas claves sobre los orígenes y la evolución del Universo. Este encuentro organizado por el Instituto de Física Corpuscular de Valencia (IFIC), un centro del Consejo Superior de Investigaciones Científicas y la Universitat de València que participa en el provecto, servirá para redactar el informe definitivo que se presentará a Bruselas con el fin de obtener financiación para una instalación de investigación básica cuvo coste rondará los 600 millones de euros, según informó aver el director del programa, Marcello Giorgi, del Istituto Nazionale di Fisica Nucleare de Italia.

Este proyecto internacional conocido como SuperB, que despegará definitivamente en el Cap i Casal tras dos años de reuniones científicos en París, Roma y San Francisco, se basa en la puesta en marcha de un supercolisionador de electrones y sus antipartículas, los positrones-que presentan las mismas características que el electrón pero con carga positiva-... que «ayudará a entender porqué en el Universo predomina la materia sobre la antimateria, un enigma no resuelto por el modelo estándar de física de partículas», explica el científico del IFIC Miguel Angel Sanchis, uno de los organizadores del encuentro.

#### Materia y antimateria

Giorgi relata que el Big Bang, la gran explosión de la que surgió el Universo, «en teoría debería de haber producido la misma cantidad de materia que antimateria, sin embargo la observación astrofísica no ha hallado ni rastro de dichas antipartículas», que son como un reflejo en el espejo de las partículas. «Es decir, con la misma masa pero distinta carga eléctrica», apunta Sanchis.



DETECTOR DE PARTÍCULAS. Una imagen del detector de partículas del del SLAC de California, que será muy similar al del nuevo supercolisionador de Roma.

#### FÍSICA DE ALTAS ENERGÍAS

#### La carrera por resolver las claves del Universo despega en Valencia

Científicos europeos y de EE UU cierran en el Cap i Casal el diseño de un nuevo **supercolisionador de electrones y positrones** único en el mundo

Fernando Martínez, uno de los cinco científicos del IFIC que lleva 10 años investigando en el colisionador de electrones y positrones del SLAC (Siglas en inglés de Stanford Linear Accelerator Center) de California, detalla que el SuperB es como una fábrica en donde se producirán «partículas y antipartículas de forma masiva, lo

 La instalación de investigación básica costará 600 millones de euros cual permitirá estudiar sus propiedades con una precisión e intensidad sin precedentes»

Martínez narra que al acelerar electrones a casi la velocidad de la luz y lanzarles positrones en sentido contrario se produce una colisión en la que partículas y antipartículas «se aniquilan». «El producto de ese choque genera nuevas partículas (como es el caso de los quarks, que son parte constituyente de los protones y los neutrones del núcleo del átomo), que se desintegran casi al instante, en apenas 1,6 picosegundos o lo que es lo mismo, un segundo precedido de 12 ceros».

El SuperB incrementará en cien veces el número de colisiones que genera el SLAC, con lo que se podrán analizar con un detalle desconocido hasta hora la desintegración de esas partículas y antipartículas de vida infinitesimalmente corta. Sanchis concluve que el super colisionador complementará al LHC, el gran acelerador de partículas del Laboratorio Europeo de Física de Partículas (CERN) de Suiza, «Son como un telescopio y un microscopio, va que la precisión del SuperB permitirá estudiar modos de desintegración, cuya observación no es posible en el LHC».



PROYECTO «SUPERB» CAMPUS DE LA UNIV. ROMA-TOR VERGATA

#### Bajo un complejo de Calatrava

Marcello Giorgi, director del provecto «SuperB», adelantó aver que el anillo subterráneo del supercolisionador ---cuvo perímetro mide 3 kilómetros- se excavará en el subsuelo del gran complejo de piscinas deportivas que levanta el arquitecto valenciano Santiago Calatrava en el campus de la Universidad de Roma-Tor Vergata para el Mundial de Natación de 2009. El anillo y la galería de un kilómetro que lo completa serán perforados por la tuneladora que está ampliando la red de metro de la Ciudad Eterna hasta dicho campus.

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## SuperB cost and governance

- SuperB will proceed as a "regional initiative", in line with the CERN Council Strategy group recommendation
- Total cost under 500 M€
  - Approx. 350 M€ needed as new money
- Governance similar to XFEL & FAIR
  - International committee formed by the interested funding agencies

#### CDR includes a cost estimate

Costs are presented "ILC-style", with replacement value for reusable PEP-II/BABAR components



Possible savings from reusing other hardware not yet considered in detail

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## Summary

- The case for flavour physics in the LHC era is compelling
  - strong complementarity with energy frontier
  - requires peak luminosity L<sub>peak</sub> >10<sup>36</sup>/cm<sup>2</sup>/s
- SuperB is the ideal tool to explore the new phenomenology
  - based on a radically new accelerator concept
- Strong European initiative to probe this window on new physics
  - explore the flavour treasure chest by mid-2010s
  - expect further developments within 6-9 months 47

#### CDR includes a cost estimate

		EDIA	Labor	M\&S	Rep.Val.
WBS	Item	mm	тт	kEuro	kEuro
1	Accelerator	5429	3497	191166	126330
1.1	Project management	2112	96	1800	0
1.2	Magnet and support system	666	1199	28965	25380
1.3	Vacuum system	620	520	27600	14200
1.4	RF system	272	304	22300	60000
1.5	Interaction region	370	478	10950	0
1.6	Controls, Diagnostics, Feedback	963	648	12951	8750
1.7	Injection and transport systems	426	252	86600	18000

		EDIA	Labor	M\&S	Rep.Val.
WBS	Item	mm	mm	kEuro	kEuro
2.0	Site	1424	1660	105700	0
2.1	Site Utilities	820	1040	31700	0
2.2	Tunnel and Support Buildings	604	620	74000	0

#### CDR includes a cost estimate

		EDIA	Labor	M\&S	Rep.Val.
WBS	Item	mm	mm	kEuro	kEuro
1	SuperB detector	3391	1873	40747	46471
1.0	Interaction region	10	4	210	0
1.1	Tracker (SVT + L0 MAPS)	248	348	5615	0
1.1.1	SVT	142	317	4380	0
1.1.2	L0 Striplet option	23	33	324	0
1.1.3	L0 MAPS option	106	32	1235	0
1.2	DCH	113	104	2862	0
1.3	PID (DIRC Pixilated PMTs + TOF)	110	222	7953	6728
1.3.1	DIRC barrel - Pixilated PMTs	78	152	4527	6728
1.3.1	DIRC barrel - Focusing DIRC	<i>92</i>	179	6959	6728
1.3.2	Forward TOF	32	70	3426	0
1.4	EMC	136	222	10095	30120
1.4.1	Barrel EMC	20	5	171	30120
1.4.2	Forward EMC	73	152	6828	0
1.4.3	Backward EMC	42	65	3096	0
1.5	IFR (scintillator)	56	54	1268	0
1.6	Magnet	87	47	1545	9623
1.7	Electronics	286	213	5565	0
1.8	Online computing	1272	34	1624	0
1.9	Installation and integration	353	624	3830	0
1.A	Project Management	720	0	180	0

NB. Items in italics (L0 striplet, focusing DIRC) are not included in the baseline

#### CDR includes a schedule

- Impossible to read here, check the CDR
- Includes site construction, PEP-II & BaBar disassembly, shipping, reassembly, etc.
- Five years from T0 to commissioning



Figure 5-1. Overall schedule for the construction of the SuperB project.

Unit	$\mathrm{Super}B$	$\mathrm{Super}B$	ILC
	LER	HER	DRs
Beam energy (GeV)	4	7	5
Circumference (m)	2249	2249	6695
Particles per bunch	$6.16\times10^{10}$	$3.52\times 10^{10}$	$2  imes 10^{10}$
Number of bunches	1733	1733	2767
Average current (A)	2.28	1.30	0.40
Horizontal emittance (nm)	1.6	1.6	0.8
Vertical emittance (pm)	4	4	2
Bunch length (mm)	6	6	9
Energy spread (%)	0.084	0.09	0.13
Momentum compaction	$1.8  imes 10^{-4}$	$3.1  imes 10^{-4}$	$4.2\times 10^{-4}$
Transverse damping time (ms)	32	32	25
RF voltage (MV)	6	18	24
RF frequency (MHz)	476	476	650

**Table 3-2.** Comparison between parameters for the SuperB storage rings and the ILC damping rings.

#### Compare to ILC "value estimate"

Costs are presented "ILC-style", with replacement value for reusable PEP-II/ BABAR components

	Tota	ls	337,613 k€	€ 172,801 k€
Detector	283	156	40,747	46,471
Site	119	138	105,700	0
Accelerator	452	291	191,166	126,330
	EDIA [my]	Labor [my]	M & S [k€]	Replacement value [k€]

SHARED VALUE =	4.87 Billion ILC VALUE UNITS	NB II C costs do not
SITE-DEPENDENT VALUE =	1.78 Billion ILC VALUE UNITS	include detector, land
TOTAL VALUE =	6.65 Billion ILC VALUE UNITS	
(shared + site-dependent)	= 5,519,500 k€	

LABOUR =	22 million person-hours = 13,000 person-years
	(assuming 1700 person-hours per person-year)

1 ILC VALUE UNIT =

1 US Dollar (2007) = 0.83 Euros = 117 Yen

MORE THAN AN ORDER OF MAGNITUDE DIFFERENCE!

- 320 Signatures
- About 85 institutions
- 174 Babar members
- 65 non Babar experimentalists.





#### Signatures breakdown by country

Signatures breakdown by type

Couplings and Scales  

$$L = L_{SM} + \sum_{k=1} \left( \sum_{i} c_{i}^{k} Q_{i}^{(k+4)} \right) / \Lambda^{k}$$

- New physics effects are governed by:
  - new physics scale  $\Lambda$
  - effective flavour-violating couplings  $c_i$ 
    - couplings may have a particular pattern (symmetries)
    - coupling strengths can vary (different interactions)
- If  $\Lambda$  known from LHC, measure  $c_i$
- If  $\Lambda$  not known, measure  $c_i / \Lambda$

## The Worst Case Scenario

- Can new physics be flavour blind?
  - No, it must couple to Standard Model, which violates flavour
- What is the minimal flavour violation?
  - new physics follows Standard Model pattern of flavour and CP violation

G. D'Ambrosio, G.F. Giudice, G. Isidori, A. Strumia, NPB 645, 155 (2002)

 even in this unfavourable scenario <u>SuperB is</u> <u>still sensitive</u>, up to new physics particle masses of 600-1000 GeV

(analysis relies on CKM fits and improvements in lattice calculations)

MFV is a long way from being verified!

### MFV Confronts the Data

- Current experimental situation
  - some new physics flavour couplings are small

Minimal flavour violation

all new physics flavour couplings are zero

MFV is a long way from being verified! Need to establish correlations between different flavour sectors (B<sub>d</sub>,B<sub>s</sub>,K)

## New Physics Sensitivity in MFV

$$\begin{aligned} \mathcal{H}_{\text{eff}}^{\Delta F=2} &= \mathcal{H}_{\text{SM}} + \mathcal{H}_{\text{NP}} = \left(V_{tq}V_{tq'}^*\right)^2 \left(\frac{S_0(x_t)}{\Lambda_0^2} + \frac{a_{\text{NP}}}{\Lambda^2}\right) (\bar{q}'q)_{(V-A)} (\bar{q}'q)_{(V-A)} \\ S_0(x_t) &\rightarrow S_0(x_t) + \delta S_0, \quad |\delta S_0| = O\left(4\frac{\Lambda_0^2}{\Lambda^2}\right), \quad \Lambda_0 = \frac{\pi Y_t}{\sqrt{2}G_F M_W} \sim 2.4 \text{ TeV} \\ \hline \text{Today} \\ \Lambda(\text{MFV}) &> 2.3\Lambda_0 \text{ @95C.L.} \\ \text{NP masses >200GeV} \qquad \qquad \text{NP masses >600GeV} \end{aligned}$$

- analysis relies on CKM fits and improvements in lattice calculations
- only  $\Delta F=2$  (mixing) operators considered
- further improvements possible including also  $\Delta F=1$  (especially  $b \rightarrow s\gamma$ )

## **Better Scenarios**

- Move slightly away from the worst case scenario
  - minimal flavour violation with large tan  $\boldsymbol{\beta}$ 
    - <u>SuperB sensitive</u> to scales of <u>few TeV</u>
  - next-to-minimal flavour violation
    - <u>SuperB sensitive</u> to scales <u>above 10 TeV</u>
  - generic flavour violation
    - <u>SuperB sensitive</u> to scales up to  $\sim 1000 \text{ TeV}$
- Look now at a few specific channels

#### MSSM + Generic Squark Mass Matrices

Today's central values with SuperB precision

 $\Delta m_d$ 

A<sub>SI</sub>

ΔΠ

magenta

green

cyan

blue



Real vs. imaginary parts of mass-insertion parameter  $(\delta_{13})_{11}$ 

### Lepton Flavour Violation

• SuperB is *much* more sensitive to LFV than LHC experiments, even for  $\tau \rightarrow \mu \mu \mu$ 

M.Roney @ Flavour in the LHC Era Workshop, CERN, March 2007

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Monte Carlo simulation of 5 $\sigma$  observation of  $\tau \rightarrow \mu\gamma$  at SuperB

## Hadronic b→s Penguins

#### Current B factory hot topic





#### Many channels can be measured with $\Delta S \sim (0.01-0.04)$

Observable	$B$ Factories (2 $ab^{-1}$ )	SuperB	
$S(\phi K^0)$	0.13	$0.02 \; (*)$	[0.030]
$S(\eta' K^0)$	0.05	0.01(*)	[0.020]
$S(K^0_{\scriptscriptstyle S}K^0_{\scriptscriptstyle S}K^0_{\scriptscriptstyle S})$	0.15	0.02 (*)	[0.037]
$S(K^0_s\pi^0)$	0.15	$0.02 \; (*)$	[0.042]
$S(\omega K_s^0)$	0.17	0.03~(*)	
$S(f_0K_s^0)$	0.12	$0.02\;(*)$	

#### (\*) theoretical limited

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N.B. This plot does not include results on f0KS from BaBar time-dependent Dalitz plot analysis of Ks $\pi$ + $\pi$ - (arXiv:0708.2097)

## **Correlations Distinguish Models**



Plots show parameter scans in four different SUSY breaking schemes: – mSUGRA

- U(2) flavour symmetry
- $-SU(5) + v_{R}$  degenerate  $-SU(5) + v_{R}$  non-degenerate

## Running at the Y(5S)

- Belle & CLEO have demonstrated potential for e  $^+e^- \to Y(5S) \to B_s^{~(*)}B_s^{~(*)}$
- Some important channels, such as  $B_s \rightarrow \gamma \gamma$ ,  $A_{SL}(B_s)$  are unique to Super*B*
- Problem: cannot resolve fast  $\Delta m_{c}$  oscillations
  - retain some sensitivity to  $\phi_s$ , since  $\Delta \Gamma_s \neq 0$

$$\Gamma_{\bar{B}_s \to f}(\Delta t) + \Gamma_{B_s \to f}(\Delta t) = \mathcal{N} \frac{e^{-|\Delta t|/\tau(B_s)}}{2\tau(B_s)} \Big[ \cosh(\frac{\Delta\Gamma_s \Delta t}{2}) - \frac{2\operatorname{Re}(\lambda_f)}{1+|\lambda_f|^2} \sinh(\frac{\Delta\Gamma_s \Delta t}{2}) \Big] \frac{1}{(1-24)} \Big] \frac{1}{(1-24)} \Big[ \cosh(\frac{\Delta\Gamma_s \Delta t}{2}) - \frac{2\operatorname{Re}(\lambda_f)}{1+|\lambda_f|^2} \cosh(\frac{\Delta\Gamma_s \Delta t}{2}) \Big] \frac{1}{(1-24)} \Big] \frac{1}{(1-24)} \Big] \frac{1}{(1-24)} \Big] \frac{1}{(1-24)} \Big[ \cosh(\frac{\Delta\Gamma_s \Delta t}{2}) - \frac{2\operatorname{Re}(\lambda_f)}{1+|\lambda_f|^2} \cosh(\frac{\Delta\Gamma_s \Delta t}{2}) \Big] \frac{1}{(1-24)} \Big] \frac{1}{(1-24)} \Big] \frac{1}{(1-24)} \Big] \frac{1}{(1-24)} \Big[ \cosh(\frac{\Delta\Gamma_s \Delta t}{2}) - \frac{2\operatorname{Re}(\lambda_f)}{1+|\lambda_f|^2} \sinh(\frac{\Delta\Gamma_s \Delta t}{2}) \Big] \frac{1}{(1-24)} \frac{1}{(1-24)} \Big] \frac{1}{(1-24)} \frac{1}{(1-24)} \Big] \frac{1}{(1-24)} \frac{1}{(1-24)} \frac{1}{(1-24)} \frac{1}{(1-24)} \frac{1$$

cf. D0 untagged measurement of  $\phi_{_{s}}$   $_{_{63}}$